

## What's Eating Ozone? Thermal Reactions between SO<sub>2</sub> and O<sub>3</sub> and Implications for Icy Environments

Mark Loeffler, R. L. Hudson,  
P. A. Gerakines

NASA Goddard Space Flight Center



## Background

- Surface composition of icy objects can be altered by radiolysis/photolysis
- Laboratory studies show radiolysis/photolysis easily forms ozone (O<sub>3</sub>) in many different ices
- Few detections of O<sub>3</sub> via remote sensing
- Why the difference?
  - Could O<sub>3</sub> react with other compounds in the ice?

## Where to Start?

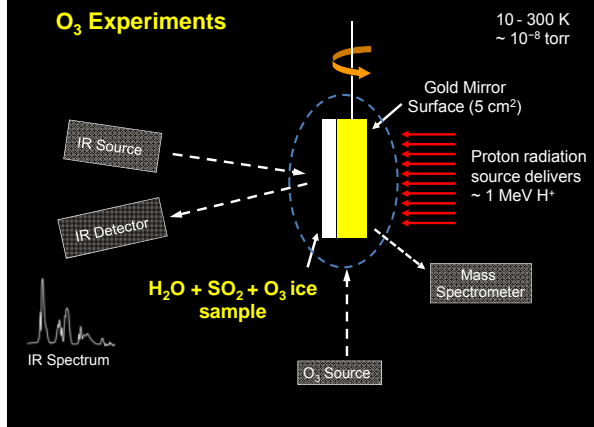
- Radiolysis prevalent on the Jovian moons
- Condensed O<sub>2</sub> has been detected on Ganymede, Callisto, Europa
  - O<sub>3</sub> only detected on Ganymede
- Sulfur compounds are present on most Jovian satellites
  - Investigate whether O<sub>3</sub> and SO<sub>2</sub> react in H<sub>2</sub>O-ice under relevant conditions

## Approach

- Grow H<sub>2</sub>O + SO<sub>2</sub> + O<sub>3</sub> (75-21-4) mixtures at 50 K
- Warm ices, while monitoring with IR spectroscopy
- To synthesize O<sub>3</sub>
  - Strike ~100 Torr of O<sub>2</sub> with Tesla coil
  - Trap O<sub>3</sub> in liquid N<sub>2</sub>



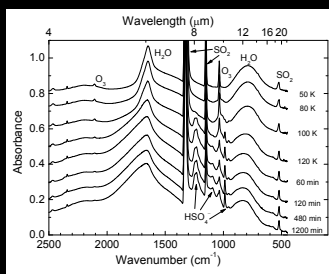
## O<sub>3</sub> Experiments



## Results

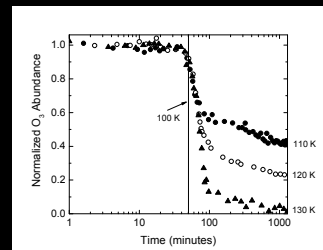
## IR Spectra During Warming

- $\text{H}_2\text{O} + \text{SO}_2 + \text{O}_3$  ice
  - Deposit at 50 K
  - Warm to 120 K (1 K / min)
- Main Sulfur Products
  - $\text{HSO}_3^-$  (80 – 100 K)
  - $\text{S}_2\text{O}_5^{2-}$  (80 – 100 K)
  - $\text{HSO}_4^-$  (> 100 K)



## Decrease in Ozone vs. Temp

- Reaction begins near 100 K
- Rate increases with temperature
- At 130 K,  $\text{O}_3$  reaches noise level in a few hours
- Temperatures are well within those observed for Jovian icy satellites



## Implications for Callisto

- Condensed  $\text{O}_2$  on trailing side but no  $\text{O}_3$
- $\text{SO}_2$  is present but more abundant on leading side
- $\text{O}_3$  on trailing side would be consumed by  $\text{SO}_2$  explaining
  - Lack of  $\text{O}_3$  on Callisto
  - Lower abundance of  $\text{SO}_2$  on trailing side

## Summary and More Implications

- $\text{O}_3$  and  $\text{SO}_2$  react readily above 100 K
  - They won't be found in same vicinity
- Callisto
  - Lack of ozone in trailing hemisphere
  - Spatial distribution of  $\text{SO}_2$
- Possibly relevant to other icy bodies
  - Mass spectrometer on Rosetta detected  $\text{O}_2$  in comet coma but no  $\text{O}_3$
  - No  $\text{O}_3$  detected in ISM and lower than expected abundance of identifiable sulfur compounds
- Future work will focus on whether other compounds similarly reactive with ozone

## Acknowledgments

- Financial support from NASA programs

NASA Solar System Workings  
Outer Planets Research  
NASA Astrobiology Institute

<http://science.gsfc.nasa.gov/691/cosmicice/>