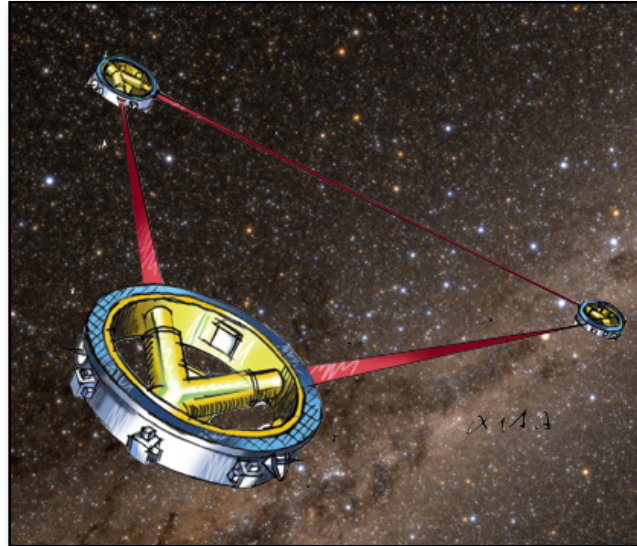


# LISA and the roles for the US and NASA

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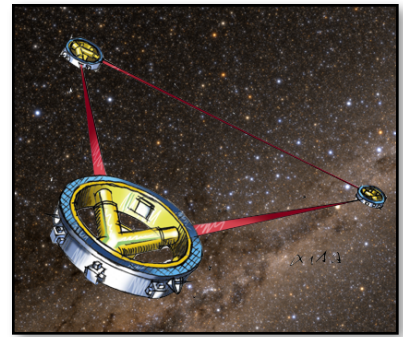
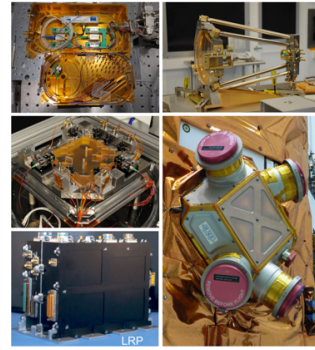
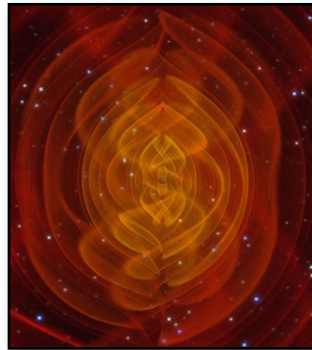
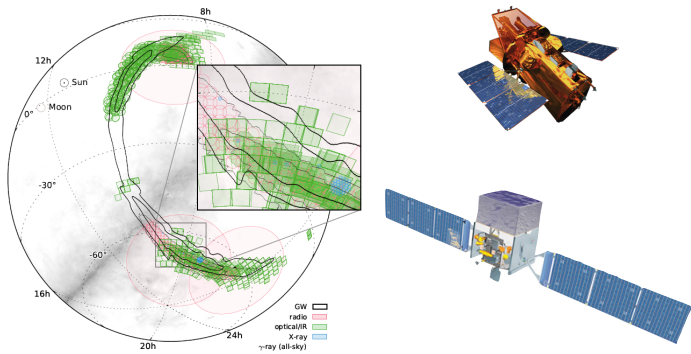
Ira Thorpe

Gravitational Astrophysics Lab (663)

NASA/GSFC

January 27<sup>th</sup>, 2017

# NASA's role in the New Astronomy



Ira Thorpe  
Gravitational Astrophysics Lab (663)

NASA/GSFC

January 27<sup>th</sup>, 2017

# Why is NASA interested in GWs?

- **From the NASA SMD Strategic Plan**
  - Objective 1.6: “Discover how the Universe works, explore how it began and evolved, and search for life on planets around other stars.”

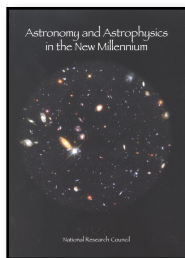


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- Jeremy Schnittmann (663)
  - “Life on Miller's Planet: The Habitability Zone Around Supermassive Black Holes”

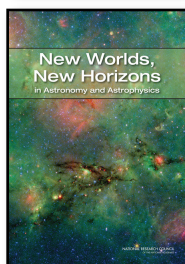


# Community Demand for GW Science



## Astrophysics “Decadal Surveys” from National Research Council

- **Astronomy & Astrophysics in the New Millennium (2001)**
  - Endorses medium-scale NASA/ESA mission for GWs



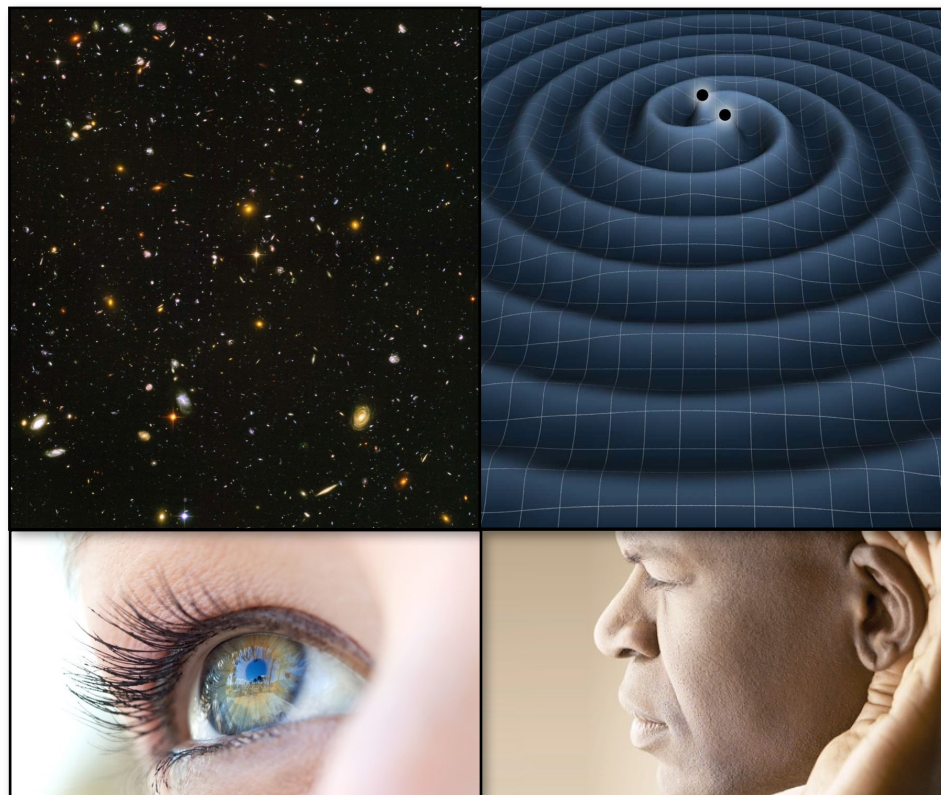
- **New Worlds, New Horizons (2010)**
  - Endorses large-scale NASA/ESA LISA mission for GWs



- **National Academies Midterm Assessment (2016)**
  - Highlights successes of LIGO and LRF
  - Recommends renewing US efforts in LISA-like mission
  - Strong endorsement for US participation in ESA-led mission

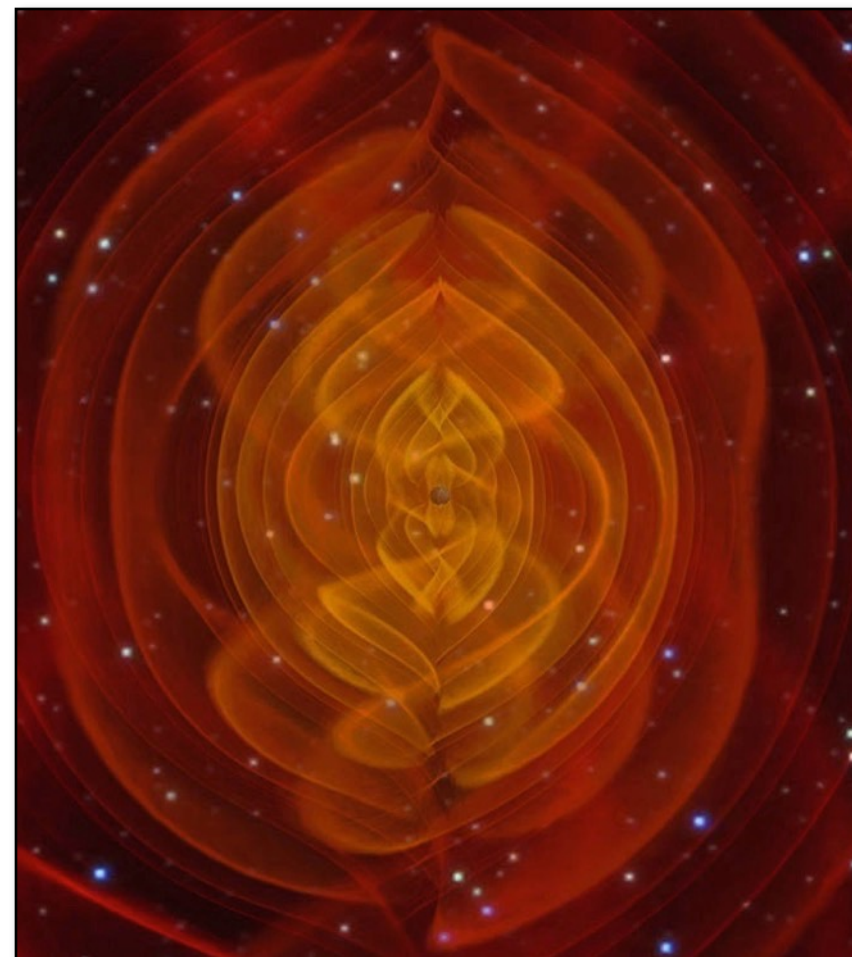
# The Promise of GW Science

- **Listening vs. Seeing**
- **Observing the Dark Sector**
- **Multi-*messenger* Astronomy**
- **Potential for discovery**



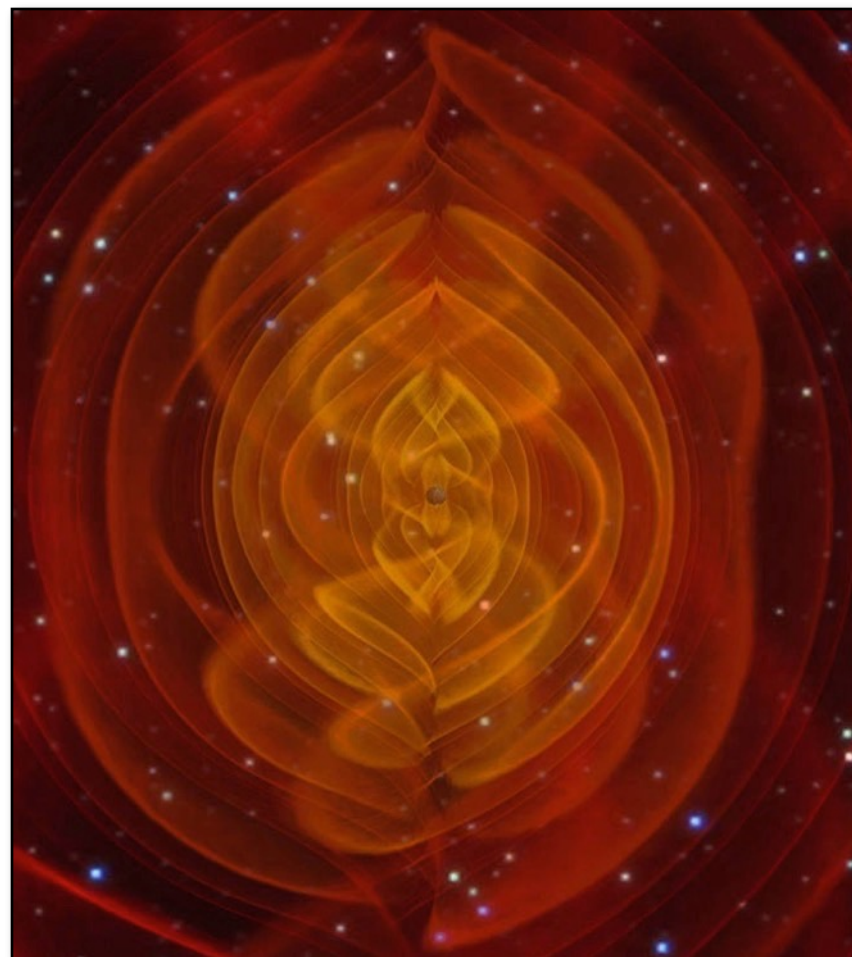
# How NASA is participating in GW science

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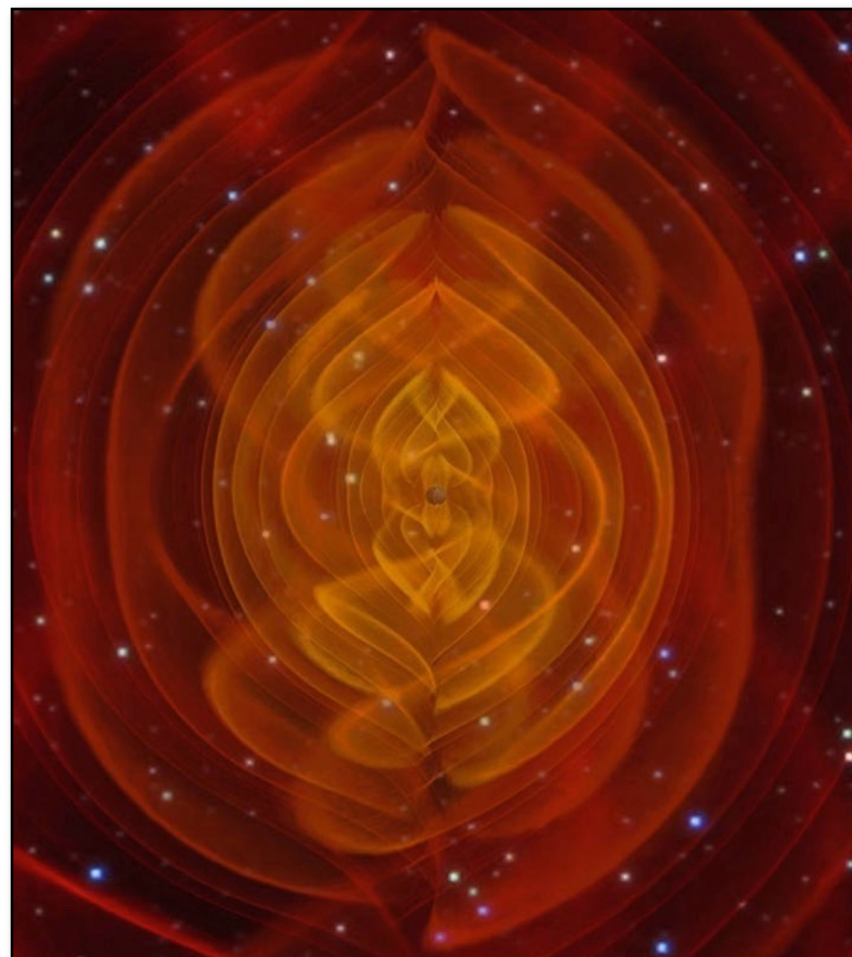
- Participating in ground-based GW astronomy, esp. EM counterparts





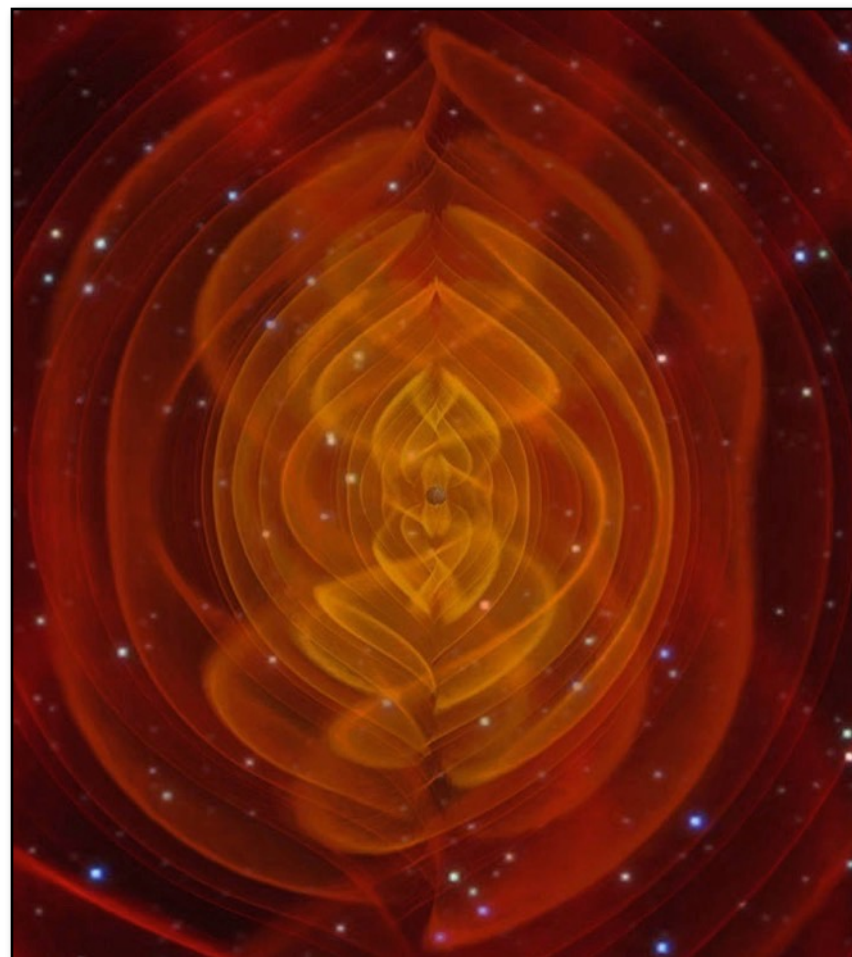
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- Participating in ground-based GW astronomy, esp. EM counterparts
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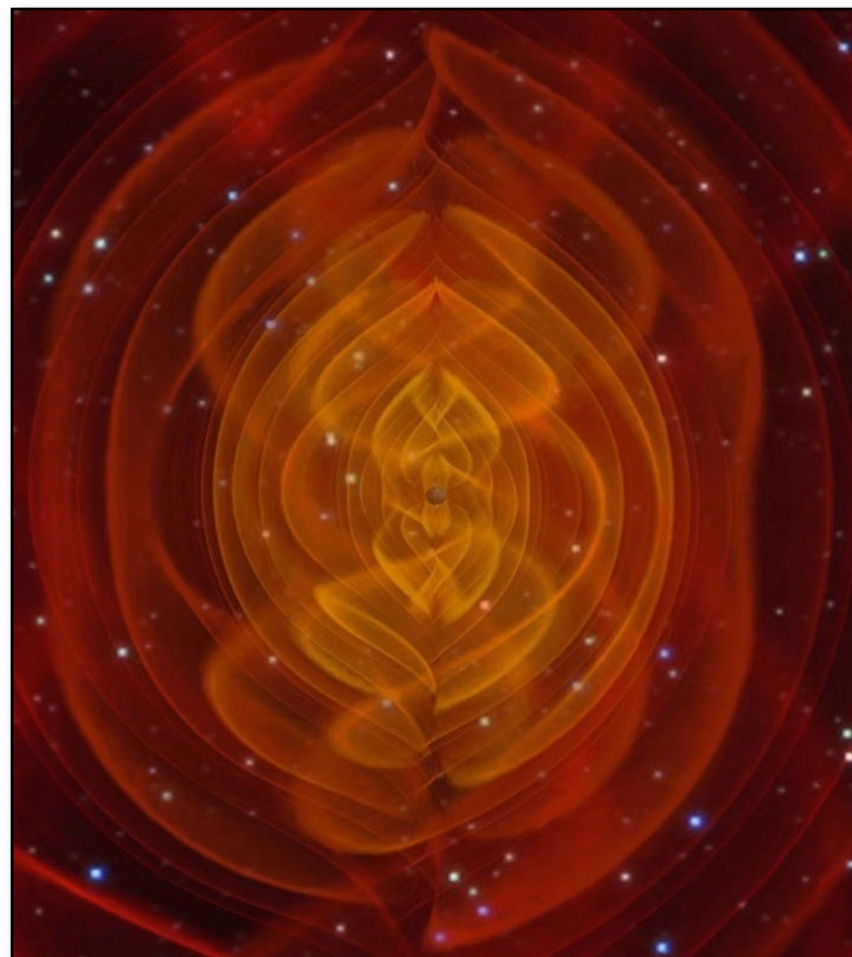
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- Developing relevant theory and analysis tools
- Participating in LISA Pathfinder operations and data analysis



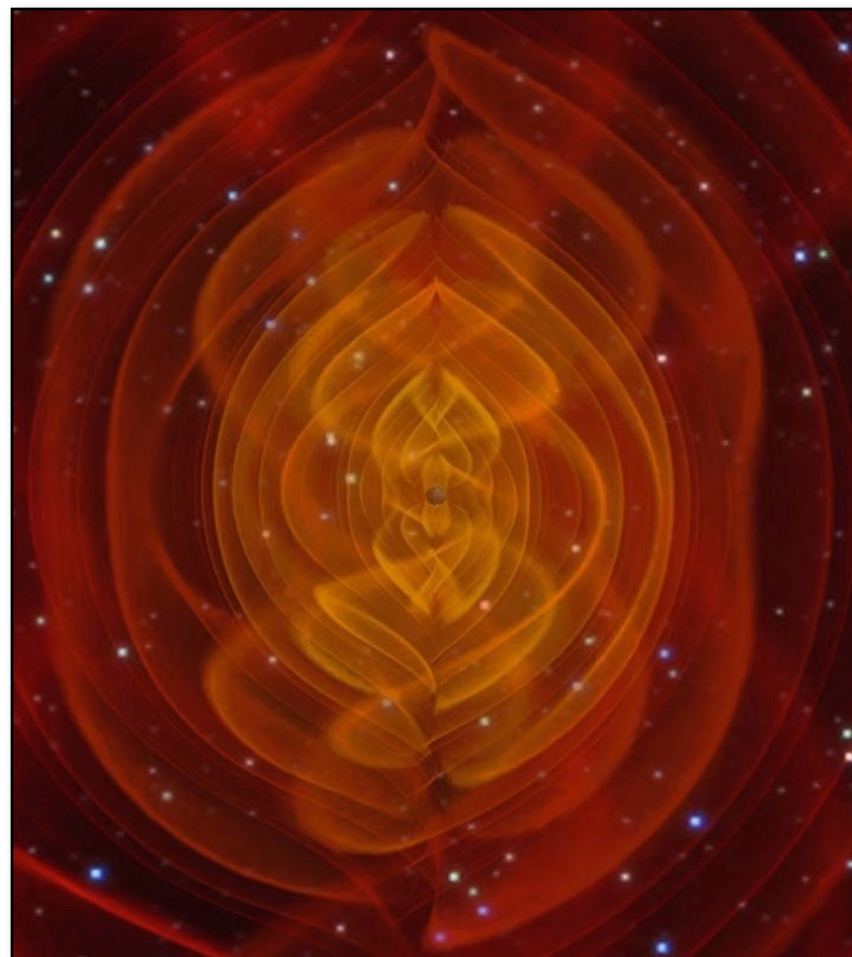
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- Developing technologies for space-based observatory



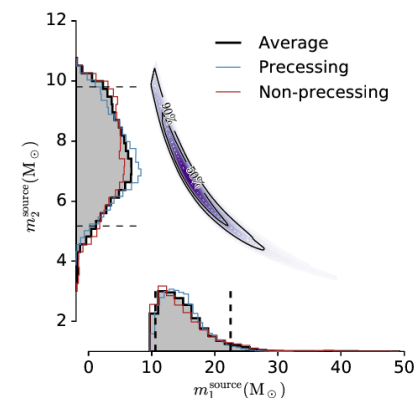
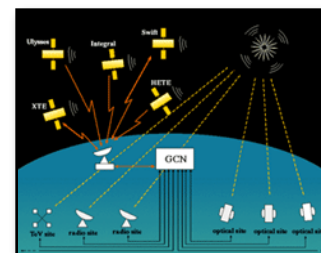
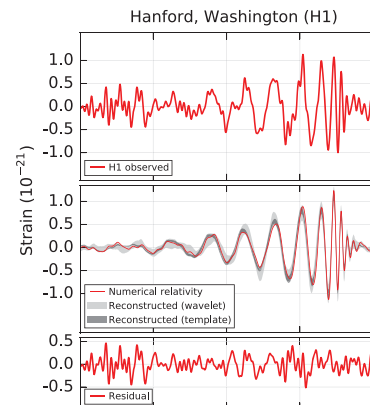
# How NASA is participating in GW science

- Participating in ground-based GW astronomy, esp. EM counterparts
- Developing relevant theory and analysis tools
- Participating in LISA Pathfinder operations and data analysis
- Developing technologies for space-based observatory
- Partnering with ESA on LISA-like mission



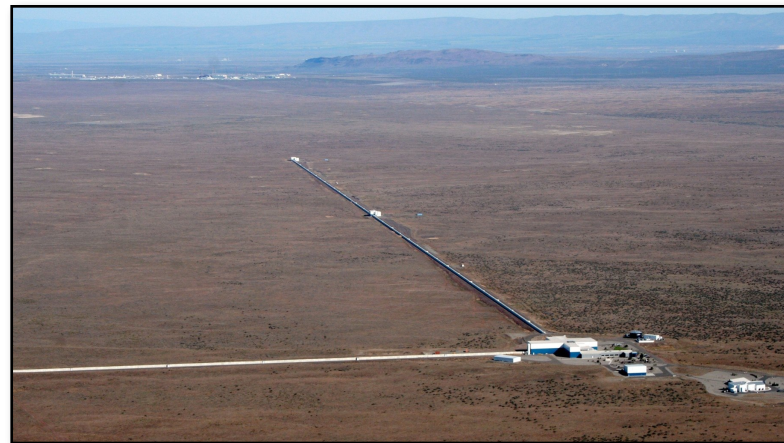
# NASA Participation in LIGO

- **LSC Members @ NASA Centers**
  - GSFC (6 members)
  - MSFC (4 members)
  - JPL/Caltech (several)
- **Activities**
  - GW detection pipelines
  - rapid source localization
  - event circulars
  - targeted follow-ups
- **Additional participation through EM follow-up**
  - Swift
  - Fermi
  - Ground-based telescopes



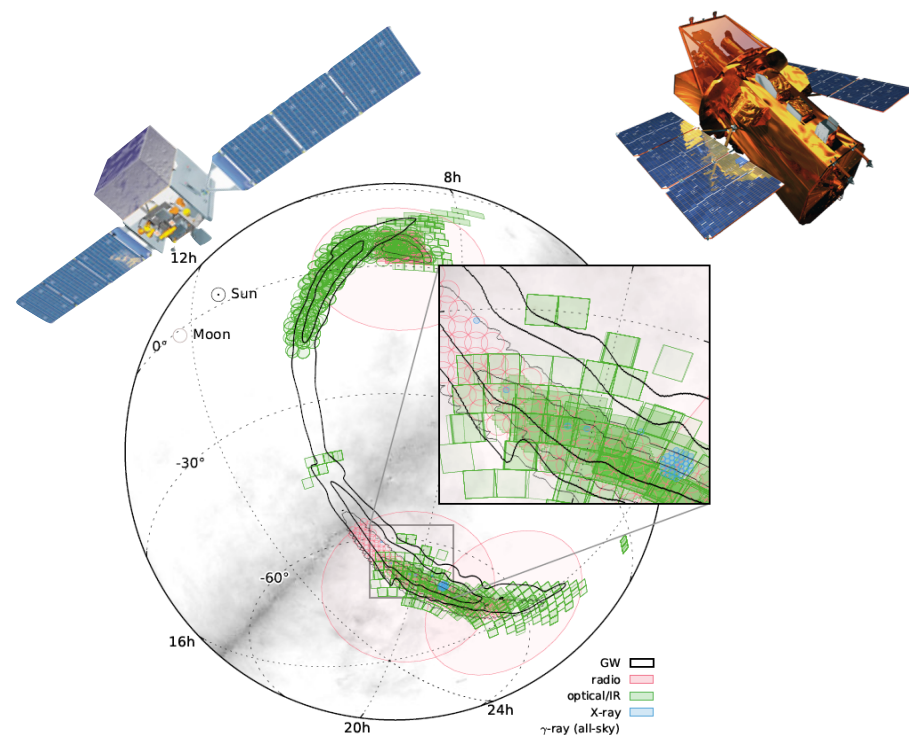
# Why follow-up is important

- **GW & EM information are complimentary**
  - bulk masses vs. atoms
  - gravity vs. temperature
  - distance vs. redshift
- **GW & EM > GW or EM**
  - source astrophysics
  - host galaxy astrophysics
  - cosmological distance ladders
  - fundamental physics tests



# NASA's Unique Role for EM follow-up

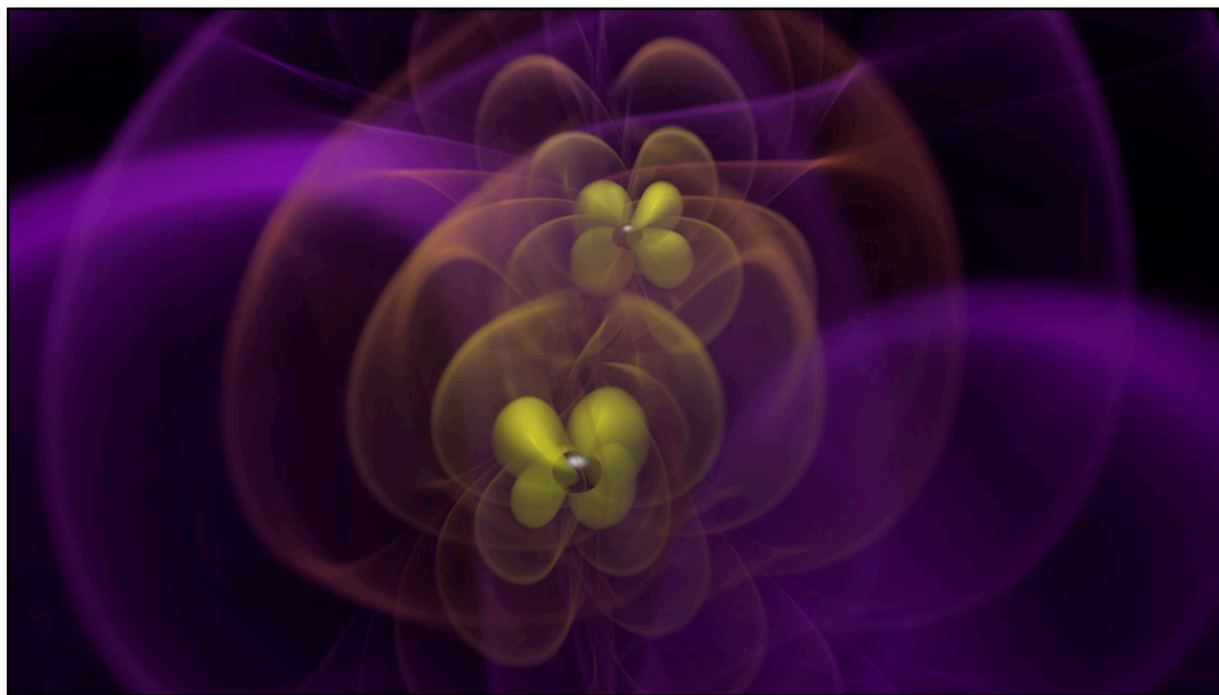
- **x-ray &  $\gamma$ -ray bands require space-based observatories**
- **Agreements in place between LVC and NASA**
  - Swift and Fermi
  - Possible Fermi Trigger from GW150914?
- **Science target for future missions**
  - Transient Astrophysics Observatory (TAO)
  - Lobster-eye optics



LIGO skymap (black contours) of GW150914 and EM follow-ups. The x-ray and gamma-ray coverage was provided by Swift and Fermi

# Theory and Data Analysis

- **Numerical relativity**
  - Solve Einstein's Equations in a supercomputer
  - Used to predict GW and EM signals from astrophysical events

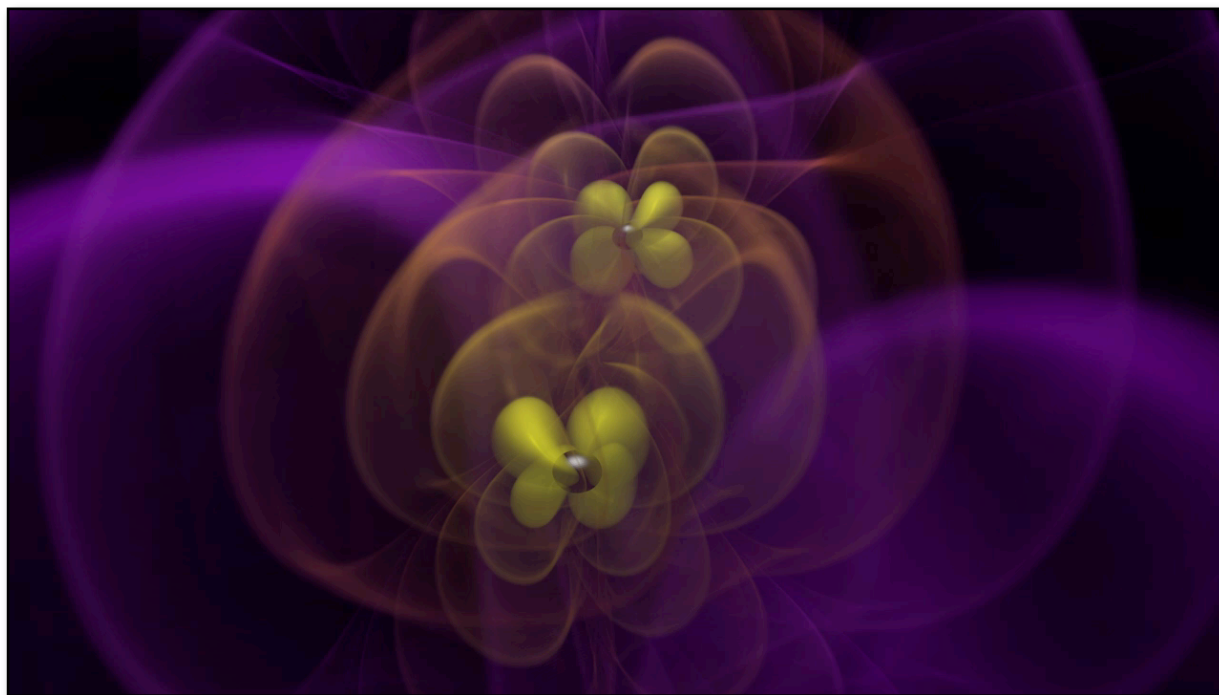


Merger of two black holes generating gravitational waves (Baker, et al.)



# Theory and Data Analysis

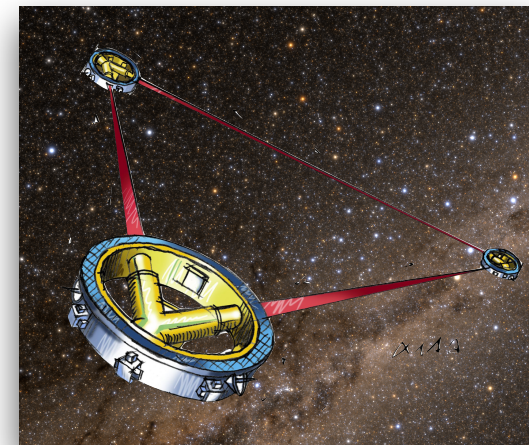
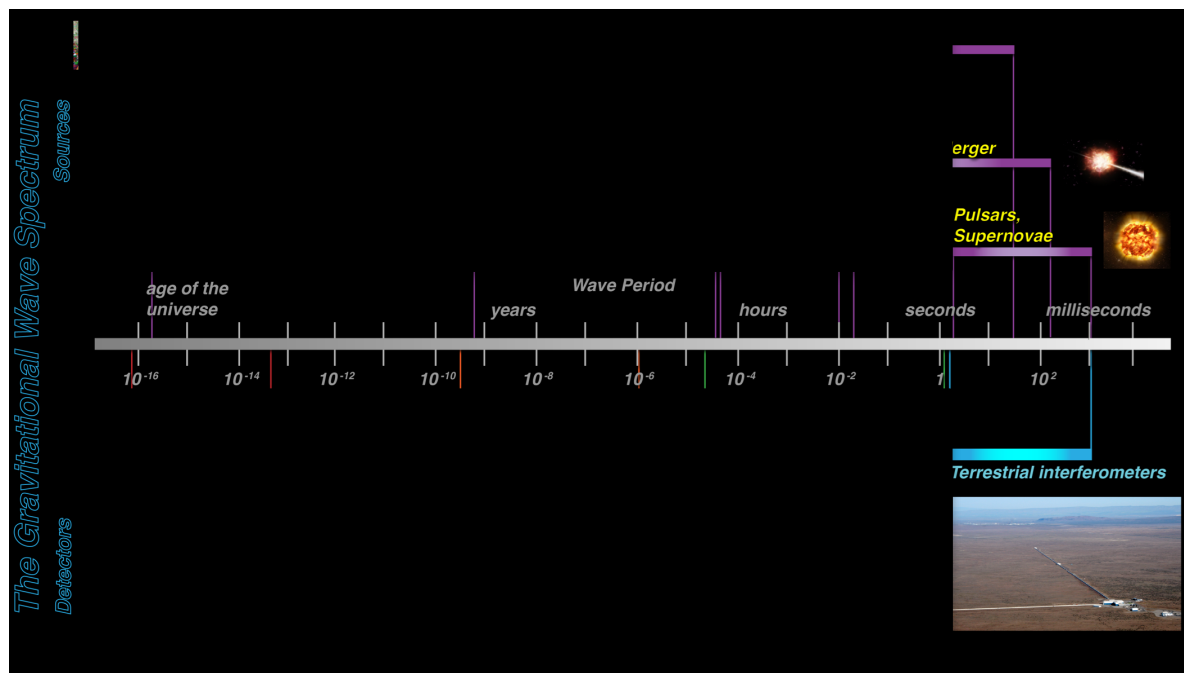
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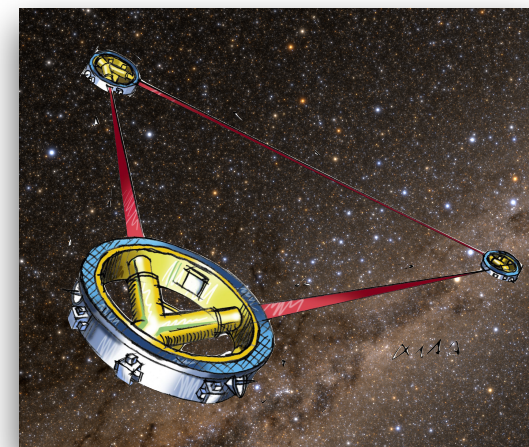
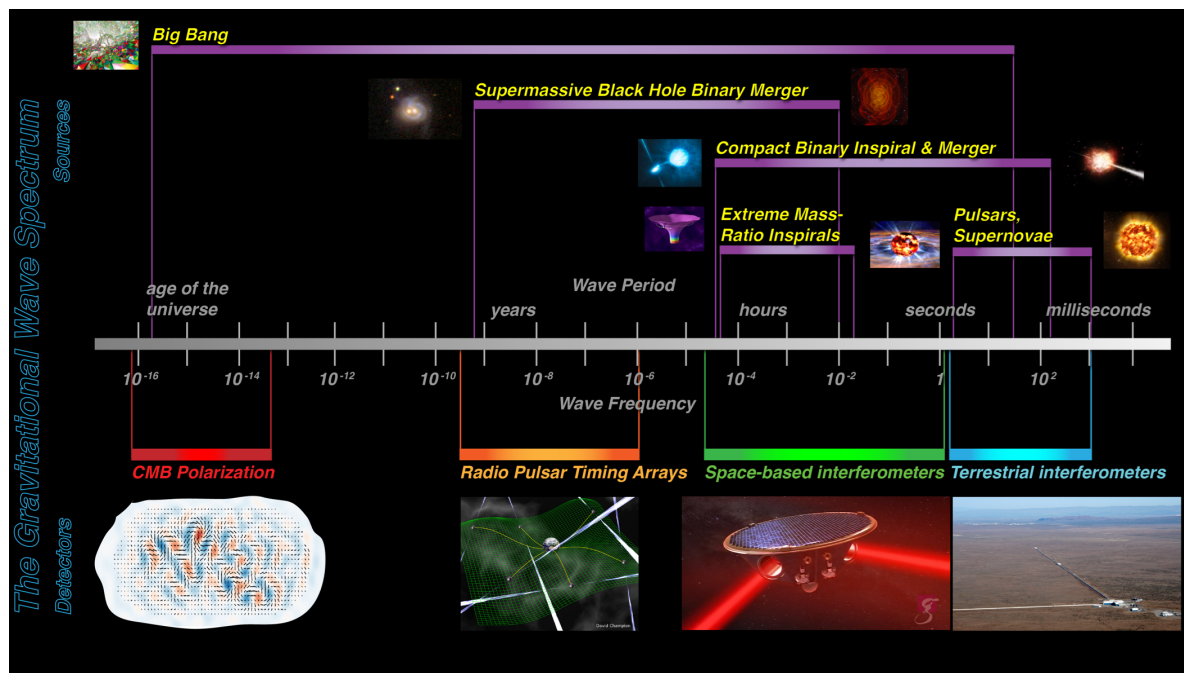
# Laser Interferometer Space Antenna

- The HST/JWST for Gravitational Waves
  - get away from terrestrial disturbances
  - access regions of the spectrum inaccessible from Earth



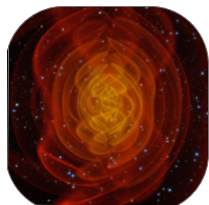
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# LISA Science

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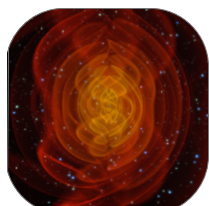


## Massive Black Hole Mergers

*visible to edge of observable universe*  
*learn about galaxy and black hole formation*  
*10 - 100 per year*

# LISA Science

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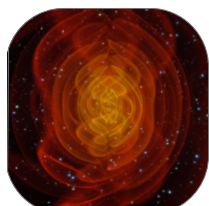
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## Extreme Mass Ratio Inspiral

*Most extreme test of General Relativity  
Learn about stellar populations in galactic cores  
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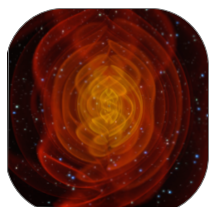
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## Close binaries in Milky Way

*Learn about compact objects and stellar evolution  
population of millions produces foreground  
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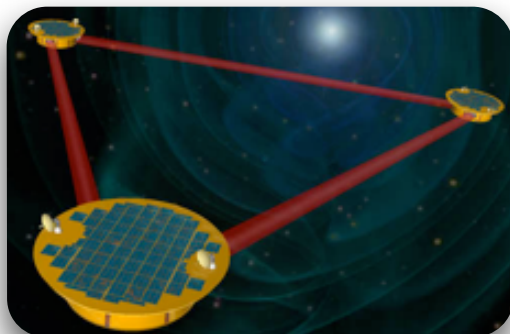
## Unexpected sources

*Exotic physics (cosmic strings, vacuum bubbles, etc.)  
Exotic astrophysics (intermediate mass black holes  
???)*

# LISA Mission Concept

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## Three-arm triangular constellation

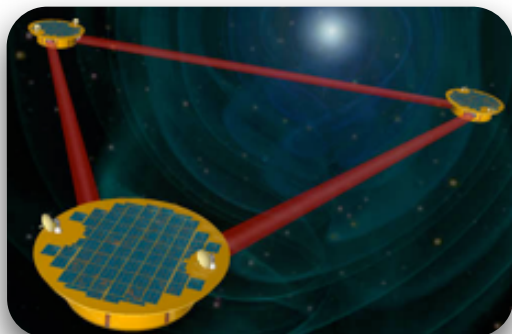


(redundancy + polarization)



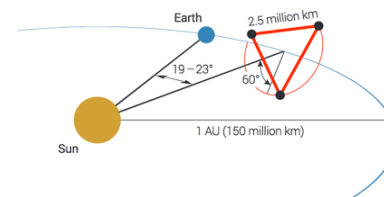
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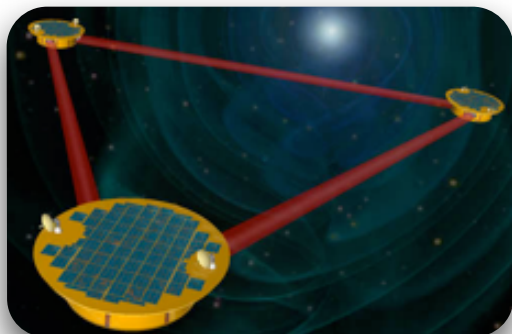
## Passively-stable, Earth-trailing heliocentric orbit



(Arm length = 2.5 Mkm)

# LISA Mission Concept

## Three-arm triangular constellation



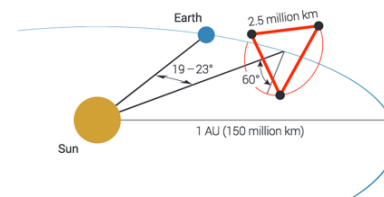
(redundancy + polarization)

## Drag-free flight to realize 'freely-falling' test mass (LPF)



$$\delta \ddot{a} \sim 3 \text{ fm/s}^2 / \sqrt{\text{Hz}}$$

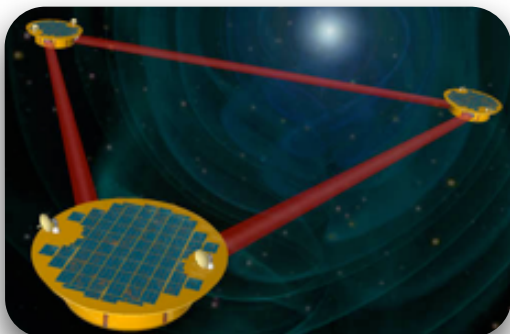
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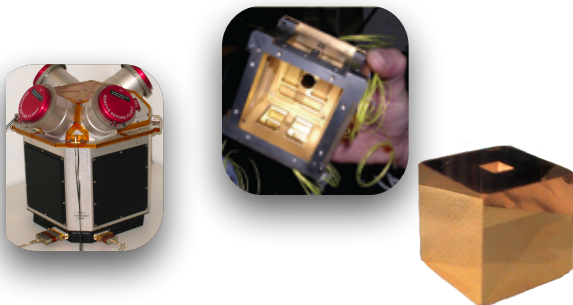
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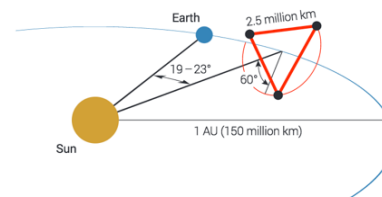
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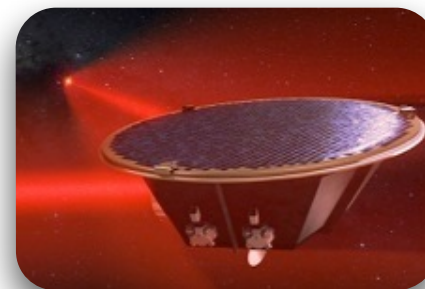
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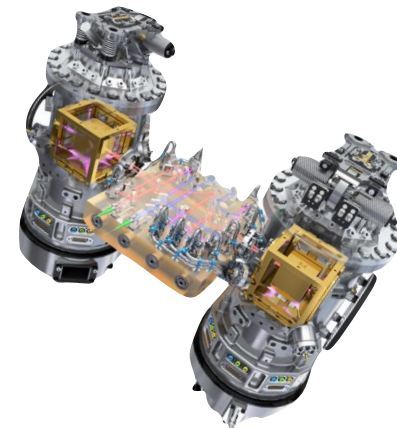
(Arm length = 2.5 Mkm)

## Heterodyne interferometry distance measurements



$$\delta \tilde{x} \sim 10 \text{ pm} / \sqrt{\text{Hz}}$$

# LISA Pathfinder



# LISA Pathfinder – DRS Operations

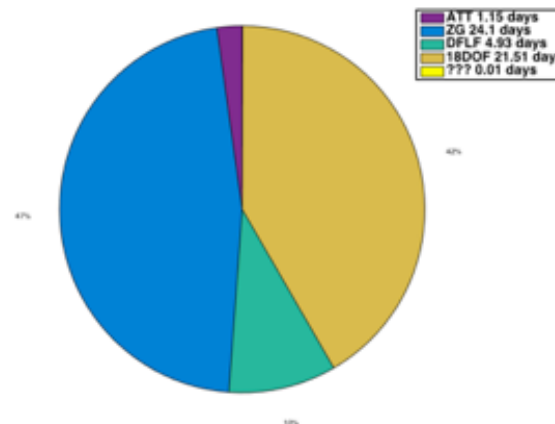
- **DRS baseline mission**

- Primary master payload on LPF from Jul - Dec, 2016
- All goals met

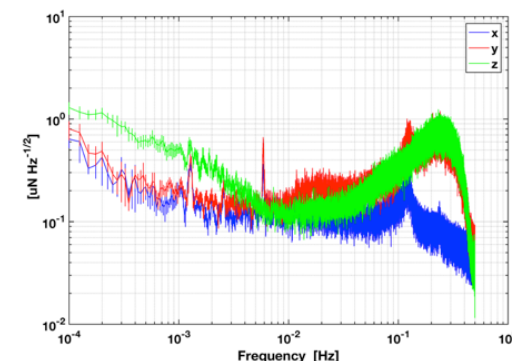
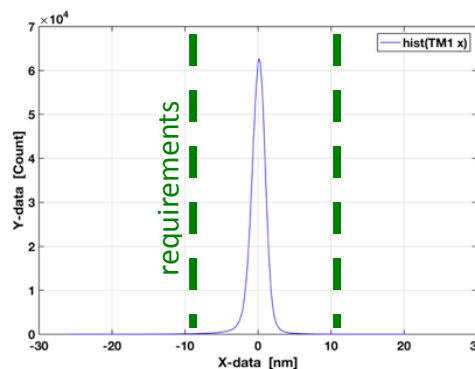
- **Extended mission**

- March 2017
- Additional thruster experiments

ST7 duty factor when in control (51.73 days)



Time-in-mode chart for DRS showing 51 days of active operations (above left). Histogram of position error during 18DOF mode, showing 2nm RMS error (above right). Thrust noise estimates showing sub-uN thruster noise performance (below)



# LISA Pathfinder – LTP Operations & Analysis

- **Small research group at GSFC**
  - Thorpe, Slutsky, & undergraduate interns
- **Participate as members of European analysis and operations consortium**
  - Work science operations shifts at ESOC in Darmstadt
  - Develop tools, run analyses, and contribute to papers

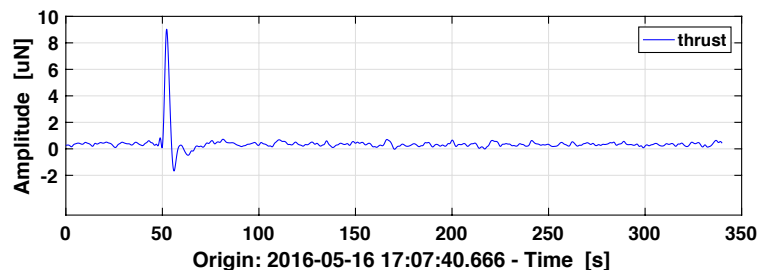
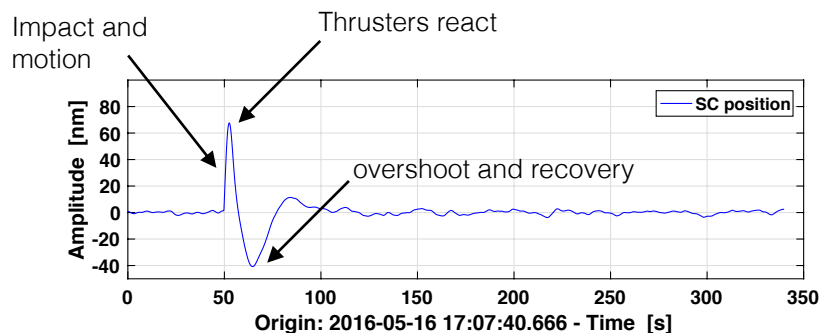
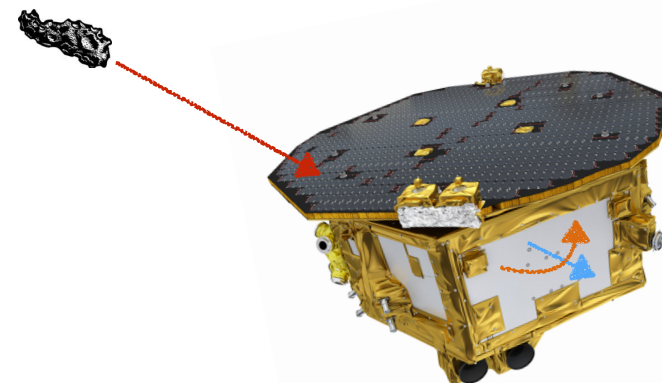


(top) Two NASA co-authors on the PRL describing LPF's preliminary results. (bottom) LTP operations at ESOC in Darmstadt, Germany.



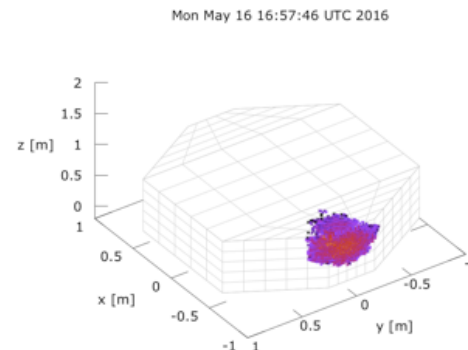
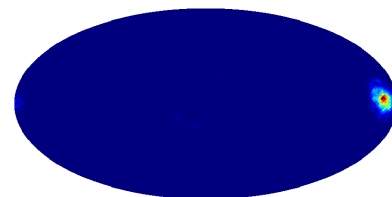
# Science of opportunity - Micrometeoroids

- Use LPF as an precision accelerometer to detect & measure impacts from dust
  - Unique tool for studying dust population
  - Potential auxiliary science for LISA



LTPDA 3.0.13.dev (R2016b), 2016-11-22 21:29:47.709 UTC, ltpda: 35dad88, iplot

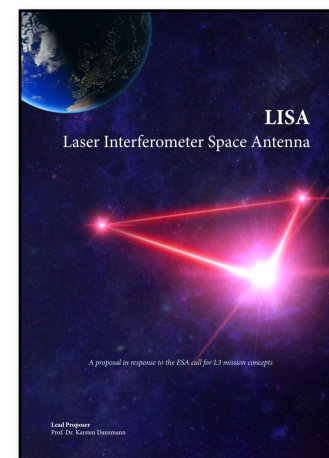
Nanogram particles impact the spacecraft and cause anomalous accelerations (above). By reconstructing the external force from the sensor signals and force commands(left) , the magnitude, direction, and location of the impact can be estimated (below)



# Current Mission Status

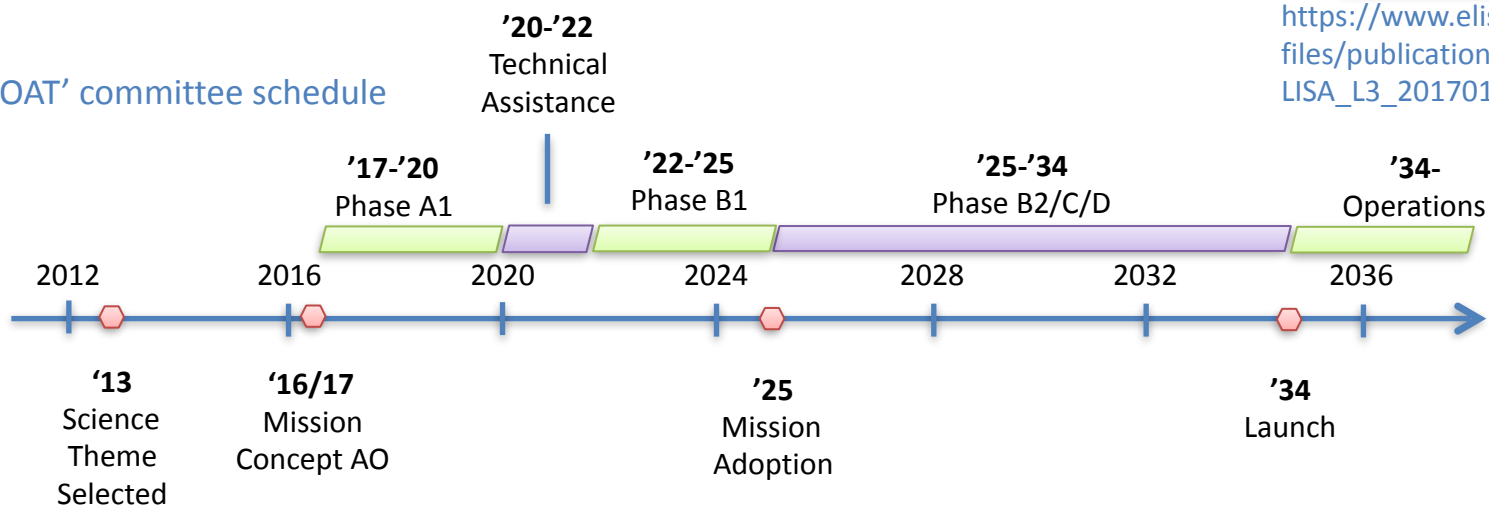
- **ESA selected GW science for 'L3' opportunity**
- **Mission concept proposal completed this month**
  - Organized by eLISA consortium (Europe)
  - Significant US participation (incl. NASA)
- **NASA/ESA Negotiations ongoing**
  - 20% Nominal contribution
  - Full participation in mission development & science

LISA L3 Mission Proposal



[https://www.elisascience.org/files/publications/LISA\\_L3\\_20170120.pdf](https://www.elisascience.org/files/publications/LISA_L3_20170120.pdf)

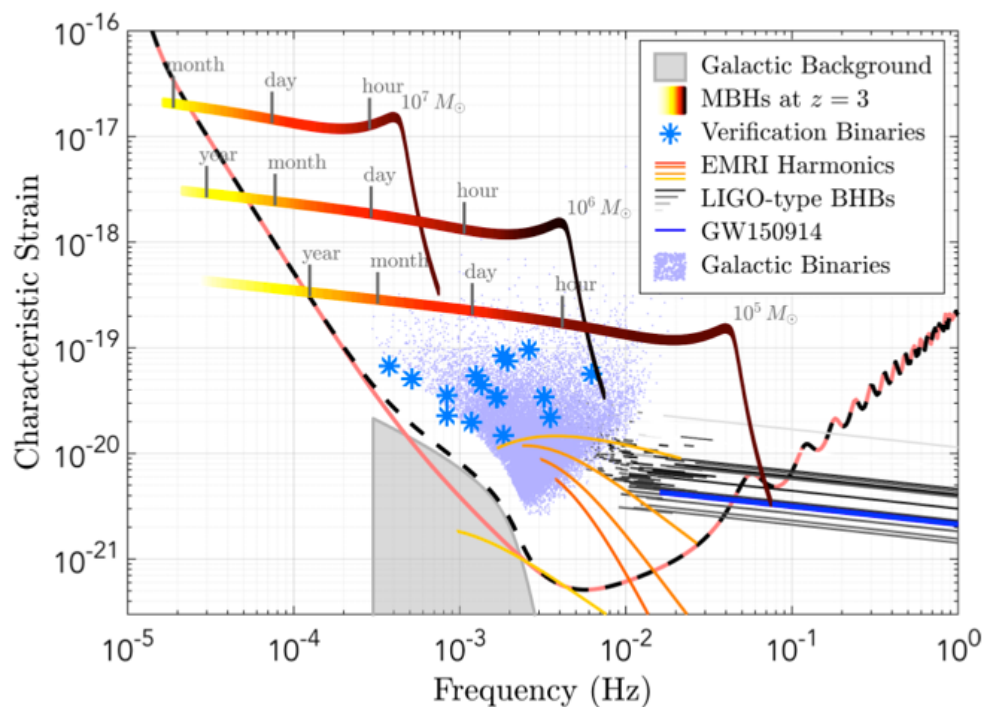
ESA 'GOAT' committee schedule





# LISA L3 Mission Concept

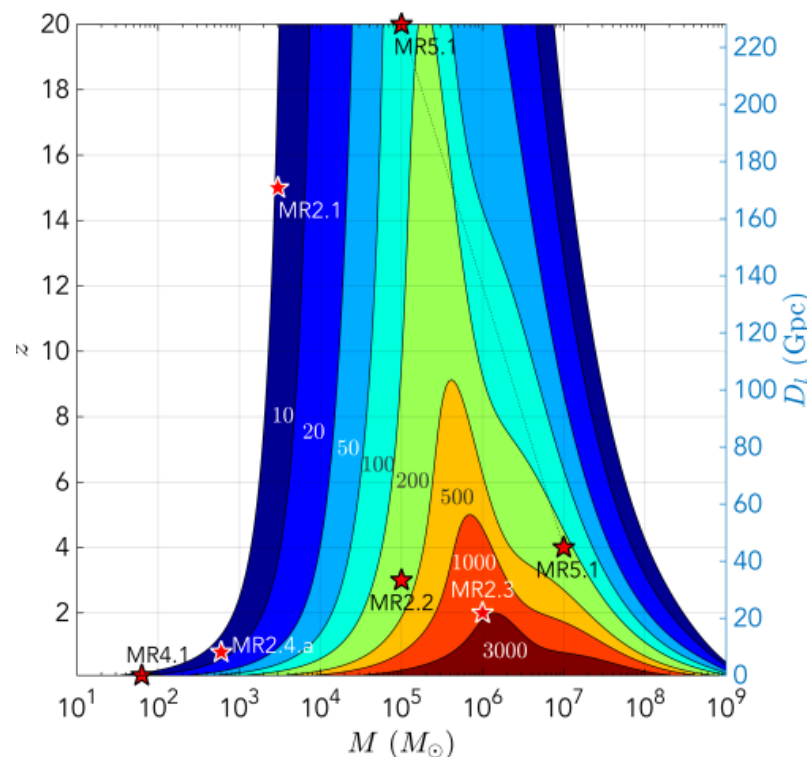
- **Adopts LISA architecture**
  - Three 2.5Mkm arms
  - 30cm telescopes
  - 2W laser (EOL)
  - LPF gravitational reference
  - 4 year nominal mission
- **Fully covers L3 science theme**
  - MBH-MBH binaries to high redshift ( $z > 20$ )
  - Extreme-mass ratio inspirals
  - Compact binaries in Milky Way
  - LIGO precursor events



LISA instrument sensitivity (black dashed line) compared with sources (colored traces) shows rich variety and number of astrophysical targets

# LISA L3 Mission Concept

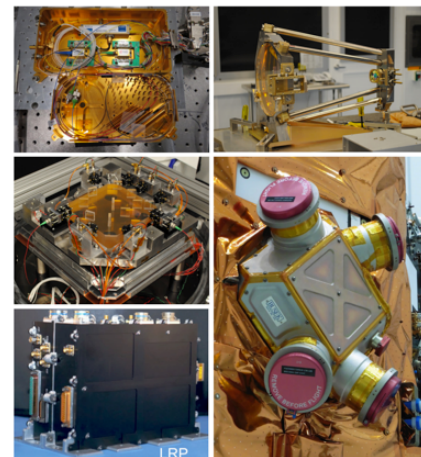
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SNR as a function of total BH mass and redshift for LISA concept for merger of black holes. LISA will conduct the **widest and deepest astrophysical survey** ever.

# How US can participate in L3

- **Hardware**
  - Major instrument systems
  - Instrument components & subsystems
  - spacecraft components
- **Mission Development**
  - Participate in trade studies
  - engineering and technical support as needed
- **Science**
  - Science case
  - Data analysis
- **Operations & Analysis**
  - Data centers
  - Science centers / guest investigator programs



US GW Technology gallery (above). NASA supercomputing resources (below)

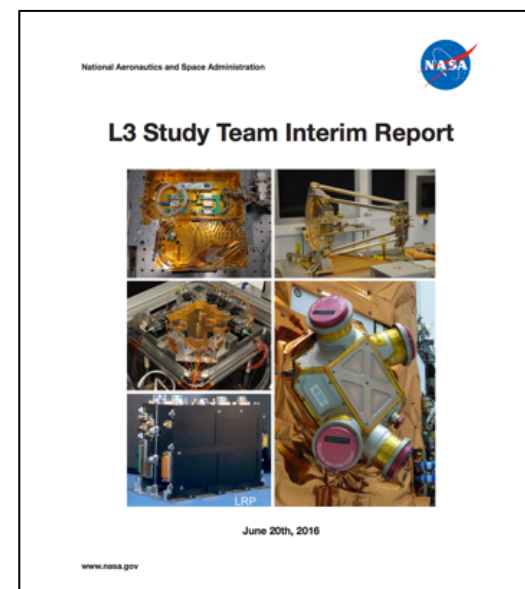


# NASA's L3 Study Team (L3ST)

- HQ-sponsored Study Team formed Dec. 2015
  - Investigate and analyze options for US participation in L3
  - Main body of 14 scientists from academia & NASA centers plus group of 6 technical experts
  - Interim Report published June 2016
  - Next Face-to-Face meeting Tuesday / Wednesday in DC



L3ST group photo, Salt Lake City, UT in April 2016

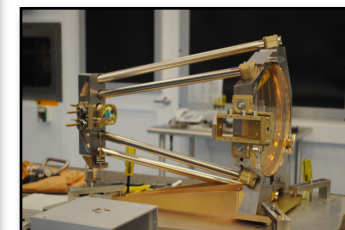
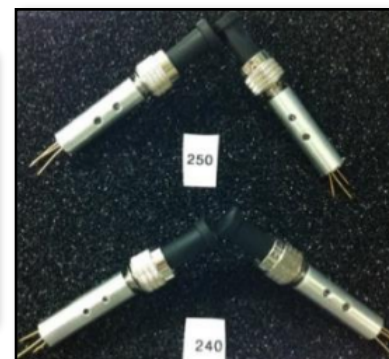
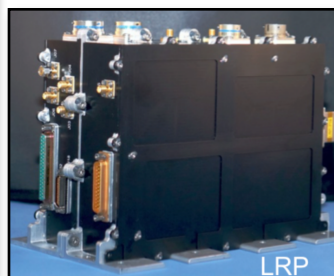
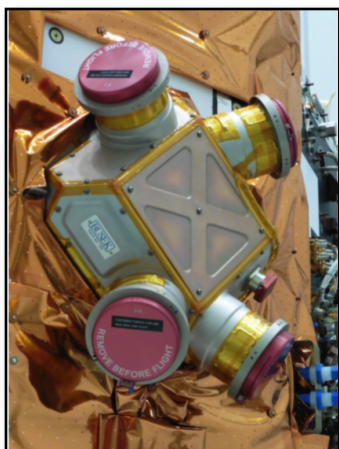
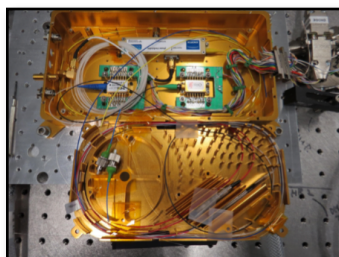
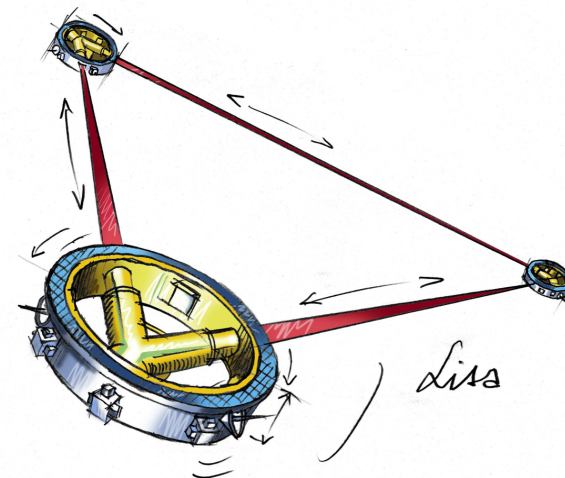


L3ST Interim Report

<https://pcos.gsfc.nasa.gov/studies/L3>

# L3 Study Office

- 'Proto-project' to coordinate US activities
  - Technology development
  - Technical interaction with ESA
  - Development of Analysis tools
  - Science Case



Potential US technology contributions (laser, thrusters, phase meter, UV discharge, telescopes)

# Summary

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- **GW Astronomy is here!**
- **NASA has important roles to play**

